Discovery of 21 μ m Sources in the Magellanic Clouds

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Spitzer Space Telescope IRS spectroscopy has discovered 8 LMC and 1 SMC carbon-rich proto-planetary nebulae, 8 of which show the unidentified 21 µm feature. The 9th object may be a very young planetary nebula. Compared to the 14 previously known Galactic 21 µm feature objects, the "PAH" bands (6.2/7.7/8.6/11.2 µm) are very strong in these spectra (compare with Hrivnak, Volk, & Kwok, 2009, ApJ, **694**, 1147 and Hrivnak, Volk, & Kwok, 2000, ApJ, **535**, 275).



The spectra are shown above with IRAC, MIPS, 2MASS, DENIS, MSX, IRAS, and UBV photometry as available. The one SMC object J004441.0-723136.44 has relatively weak 21 and 30 µm features. The red curves show model fits excluding the non-equilibrium PAH features.



Optical photometry from the OGLE project shows that the objects have low amplitude variability with periods of ~40 to ~90 days. The amplitudes are smaller and the periods shorter than for the Galactic objects.



The LMC image (IRAC [3.6]) showing the positions of the eight 21 μ m objects (blue) and the other object (green). They are found mostly in the outer regions of the LMC.

Object Name	Observed Flux (10 ⁻¹⁴ W/m ²)	Total Luminosity (L _☉)	Model T _* (K)	Model T _{dust} (r _{in}) [K]	Model V Optical Depth	21 μm feature strength	30 μm feature strength
J004441.04-723136.44	4.20	4750	6000	317	1.46	0.365	0.369
J050713.76-674848.1	8.25	6290	9500	242	13.78	0.039	0.979
J050835.91-711730.6	8.37	6550	5000	337	27.43	0.297	0.469
J051110.64-661253.7	8.55	6700	5700	253	5.63	0.5	1.55
J051228.19-690755.8	5.37	4200	30000	151	0.82		0.7
J051845.23-700534.5	8.49	6650	9000	306	3.77	0.583	0.233
J052043.86-692341.0	6.46	5060	5600	335	1.08	0.912	0.822
J052520.76-705007.5	6.43	5040	7000	261	1.52	0.735	0.281
J052902.39-661527.8	11.0	8630	6500	306	8.93	0.104	0.744

Some **object properties** are listed in the Table above. The luminosities are based on the observed fluxes plus distances of 50 and 60.1 kpc for the LMC and the SMC. Galactic 21 μ m object luminosities are estimated to be 7000-10000 L_o, subject to large distance uncertainties. Therefore these LMC/SMC luminosities are smaller than expected. Values in the right part of the Table are from the dust model fits. The 21 and 30 μ m feature strengths are given relative to the continuum. The model dust temperatures are ~100K larger than is found for Galactic objects.

A PDF version of this is available at http://www.stsci.edu/~volk.